

### PROBLEM SET 1

**P12.1)** When a molecule absorbs a photon, both the energy and momentum are conserved. If a  $\text{H}_2$  molecule at 750. K absorbs an ultraviolet photon of wavelength 225 nm, what is the change in its velocity  $\Delta v$ ? Given that its average speed is  $v_{\text{rms}} = \sqrt{3kT/m}$ , what is  $\Delta v/v_{\text{rms}}$ ?

Because momentum is conserved,

$$p_{\text{photon}} = \frac{h}{\lambda} = \frac{6.626 \times 10^{-34} \text{ J s}}{225 \times 10^{-9} \text{ m}} = 2.94 \times 10^{-27} \text{ kg m s}^{-1}$$

All this momentum is transferred to the  $\text{H}_2$  molecule

$$\Delta p_{\text{H}_2} = 2.94 \times 10^{-27} \text{ kg m s}^{-1} = m \Delta v$$

$$\Delta v_{\text{H}_2} = \frac{2.94 \times 10^{-27} \text{ kg m s}^{-1}}{2.016 \text{ amu} \times 1.661 \times 10^{-27} \text{ kg/amu}} = 0.879 \text{ m s}^{-1}$$

$$\frac{\Delta v}{v} = \frac{0.879 \text{ m s}^{-1}}{\sqrt{\frac{3RT}{M}}} = \frac{0.879 \text{ m s}^{-1}}{\sqrt{\frac{3 \times 8.314 \text{ J mol}^{-1} \text{ K}^{-1} \times 298 \text{ K}}{2.016 \times 10^{-3} \text{ kg mol}^{-1}}}} = 2.89 \times 10^{-4}$$

**P12.2)** A more accurate expression for  $\bar{E}_{\text{osc}}$  would be obtained by including additional terms in the Taylor-Mclaurin series. The Taylor-Mclaurin series expansion of  $f(x)$  in the vicinity of  $x_0$  is given by (see Math Supplement)

$$f(x) = f(x_0) + \left( \frac{df(x)}{dx} \right)_{x=x_0} (x-x_0) + \frac{1}{2!} \left( \frac{d^2 f(x)}{dx^2} \right)_{x=x_0} (x-x_0)^2 + \frac{1}{3!} \left( \frac{d^3 f(x)}{dx^3} \right)_{x=x_0} (x-x_0)^3 + \dots$$

Use this formalism to better approximate  $\bar{E}_{\text{osc}}$  by expanding  $e^{\frac{h\nu}{kT}}$  in powers of  $h\nu/kT$  out to  $(h\nu/kT)^3$  in the vicinity of  $h\nu/kT = 0$ . Calculate the relative error,  $\frac{\bar{E}_{\text{osc}} - kT}{\bar{E}_{\text{osc}}}$ , if you had not included the additional terms for  $\nu = 1.00 \times 10^{12} \text{ s}^{-1}$  at temperatures of 800., 500., and 250. K. Explain the trend you see.

The Taylor series expansion of  $e^{\frac{h\nu}{kT}}$  is  $1 + \frac{h\nu}{kT} + \frac{1}{2}\left(\frac{h\nu}{kT}\right)^2 + \frac{1}{6}\left(\frac{h\nu}{kT}\right)^3 + \dots$ . Therefore including terms up to  $\left(\frac{h\nu}{kT}\right)^3$ ,

$$\bar{E}_{osc} = \frac{h\nu}{\frac{h\nu}{kT} + \frac{1}{2}\left(\frac{h\nu}{kT}\right)^2 + \frac{1}{6}\left(\frac{h\nu}{kT}\right)^3}$$

$$\frac{E_{osc} - kT}{E_{osc}} = \left[ \frac{h\nu}{\frac{h\nu}{kT} + \frac{1}{2}\left(\frac{h\nu}{kT}\right)^2 + \frac{1}{6}\left(\frac{h\nu}{kT}\right)^3} - kT \right] \bigg/ \frac{h\nu}{\frac{h\nu}{kT} + \frac{1}{2}\left(\frac{h\nu}{kT}\right)^2 + \frac{1}{6}\left(\frac{h\nu}{kT}\right)^3}$$

$$\frac{E_{osc} - kT}{E_{osc}} = \frac{6.626 \times 10^{-34} \text{ J s} \times 1.00 \times 10^{12} \text{ s}^{-1}}{\left[ \frac{6.626 \times 10^{-34} \text{ J s} \times 1.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 800 \text{ K}} + \frac{1}{2} \left( \frac{6.626 \times 10^{-34} \text{ J s} \times 1.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 800 \text{ K}} \right)^2 + \frac{1}{6} \left( \frac{6.626 \times 10^{-34} \text{ J s} \times 1.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 800 \text{ K}} \right)^3 \right]} \left[ \frac{6.626 \times 10^{-34} \text{ J s} \times 1.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 800 \text{ K}} + \frac{1}{2} \left( \frac{6.626 \times 10^{-34} \text{ J s} \times 1.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 800 \text{ K}} \right)^2 + \frac{1}{6} \left( \frac{6.626 \times 10^{-34} \text{ J s} \times 1.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 800 \text{ K}} \right)^3 \right]} - 1.381 \times 10^{-23} \text{ J K}^{-1} \times 800 \text{ K} \times \left. \right\}$$

$$\frac{E_{osc} - kT}{E_{osc}} = -0.0306 \text{ for } 800. \text{ K. The corresponding values for } 500. \text{ K and } 250. \text{ K are } -0.0495$$

and  $-0.102$ . We expect the high temperature limit of  $\bar{E}_{osc}$  to be  $kT$ . Therefore the error is small at 800. K but becomes larger as  $T$  decreases.

**P12.7)** Show that the energy density radiated by a blackbody

$$\frac{E_{total}(T)}{V} = \int_0^\infty \rho(\nu, T) d\nu = \int_0^\infty \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{h\nu/kT} - 1} d\nu$$

depends on the temperature as  $T^4$ . (*Hint:* Make the substitution of variables  $x = h\nu/kT$ .) The definite integral  $\int_0^\infty [x^3/(e^x - 1)] dx = \pi^4/15$ . Using your result, calculate the energy density radiated by a blackbody at 1350. K and 5250. K.

$$\frac{E_{total}}{V} = \int_0^\infty \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{h\nu/kT} - 1} d\nu. \text{ Let } x = h\nu/kT; \quad dx = \frac{h}{kT} d\nu$$

$$\int_0^{\infty} \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{h\nu/kT} - 1} d\nu = \frac{8\pi k^4 T^4}{h^3 c^3} \int_0^{\infty} \frac{x^3}{e^x - 1} dx = \frac{8\pi^5 k^4 T^4}{15 h^3 c^3}$$

$$\text{At } 1350. \text{ K, } \frac{E_{\text{total}}}{V} = \frac{8\pi^5 k^4 T^4}{15 h^3 c^3} = \frac{8\pi^5 (1.381 \times 10^{-23} \text{ J K}^{-1})^4 \times (1350. \text{ K})^4}{15 \times (6.626 \times 10^{-34} \text{ J s})^3 (2.998 \times 10^8 \text{ m s}^{-1})^3} = 2.51 \times 10^{-3} \text{ J m}^{-3}$$

$$\frac{E_{\text{total}}}{V} = \frac{8\pi^5 \times (1.381 \times 10^{-23} \text{ J K}^{-1})^4 \times (5250. \text{ K})^4}{15 \times (6.626 \times 10^{-34} \text{ J s})^3 \times (2.998 \times 10^8 \text{ m s}^{-1})^3} = 0.575 \text{ J m}^{-3}$$

**P12.9)** A newly developed substance that emits 315 W of photons with a wavelength of 275 nm is mounted in a small rocket initially at rest in outer space such that all of the radiation is released in the same direction. Because momentum is conserved, the rocket will be accelerated in the opposite direction. If the total mass of the rocket is 6.75 kg, how fast will it be traveling at the end of 365 days in the absence of frictional forces?

The number of photons is given by

$$n' = \frac{E_{\text{total}}}{E_{\text{photon}}} = \frac{\text{Watts} \times \frac{1 \text{ J s}^{-1}}{\text{W}}}{\frac{hc}{\lambda}} = \frac{315 \text{ W} \times \frac{1 \text{ J s}^{-1}}{\text{W}}}{\frac{6.626 \times 10^{-34} \text{ J s} \times 2.998 \times 10^8 \text{ m s}^{-1}}{275 \times 10^{-9} \text{ m}}} = 4.36 \times 10^{20} \text{ s}^{-1}$$

The momentum of one photon is

$$p = \frac{h}{\lambda} = \frac{6.626 \times 10^{-34} \text{ J s}}{275 \times 10^{-9} \text{ m}} = 2.41 \times 10^{-27} \text{ kg m s}^{-1}$$

The force is given by the rate of change of momentum.

$$F = \frac{d(n'p)}{dt} = 2.41 \times 10^{-27} \text{ kg m s}^{-1} \times 4.36 \times 10^{20} \text{ s}^{-1} = 1.05 \times 10^{-6} \text{ kg m s}^{-2}$$

The final speed is given by

$$v = v_0 + at = \frac{F}{m}t = \frac{1.05 \times 10^{-6} \text{ kg m s}^{-2}}{6.75 \text{ kg}} \times 365 \text{ days} \times \frac{86400 \text{ s}}{\text{day}} = 4.91 \text{ m s}^{-1}$$

**P12.10)** In our discussion of blackbody radiation, the average energy of an oscillator

$E_{osc} = h\nu / (e^{h\nu/kT} - 1)$  was approximated as  $E_{osc} = h\nu / [(1 + h\nu/kT) - 1] = kT$  for  $h\nu/kT \ll 1$ . Calculate the relative error  $= (E - E_{approx})/E$  in making this approximation for  $\nu = 4.00 \times 10^{12} \text{ s}^{-1}$  at temperatures of 6000., 2000., and 500. K. Can you predict what the sign of the relative error will be without a detailed calculation?

$$\text{Relative Error} = \frac{E - E_{approx}}{E} = \frac{\left( \frac{h\nu}{\exp\left(\frac{h\nu}{kT}\right) - 1} \right) - kT}{\left( \frac{h\nu}{\exp\left(\frac{h\nu}{kT}\right) - 1} \right)}$$

$$= \frac{\left( \frac{h\nu}{\exp\left(\frac{h\nu}{kT}\right) - 1} \right) - kT}{\left( \frac{h\nu}{\exp\left(\frac{h\nu}{kT}\right) - 1} \right)}$$

$$= \frac{\left( \frac{6.626 \times 10^{-34} \text{ J s} \times 4.00 \times 10^{12} \text{ s}^{-1}}{\exp\left(\frac{6.626 \times 10^{-34} \text{ J s} \times 4.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 6000 \text{ K}}\right) - 1} \right) - 1.381 \times 10^{-23} \text{ J K}^{-1} \times 6000 \text{ K}}{\left( \frac{6.626 \times 10^{-34} \text{ J s} \times 4.00 \times 10^{12} \text{ s}^{-1}}{\exp\left(\frac{6.626 \times 10^{-34} \text{ J s} \times 4.00 \times 10^{12} \text{ s}^{-1}}{1.381 \times 10^{-23} \text{ J K}^{-1} \times 6000 \text{ K}}\right) - 1} \right)}$$

$$\text{Relative Error} = \frac{E - E_{approx}}{E} = -0.0162 \text{ for } T = 6000. \text{ K.}$$

The results for 2000. K, and 500. K are  $-0.0496$  and  $-0.219$ .

Because  $E_{approx} = kT > E$ , the relative error  $= \frac{E - E_{approx}}{E}$  is always a negative number.

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**P12.14)** The distribution in wavelengths of the light emitted from a radiating blackbody is a sensitive function of the temperature. This dependence is used to measure the temperature of hot objects, without making physical contact with those objects, in a technique called *optical pyrometry*. In the limit  $(hc/\lambda kT) \gg 1$ , the maximum in a plot of  $\rho(\lambda, T)$  versus  $\lambda$  is given by  $\lambda_{\max} = hc/5kT$ . At what wavelength does the maximum in  $\rho(\lambda, T)$  occur for  $T = 850.$ ,  $1300.$ , and  $5500.$  K?

According to Example Problem 14.1,  $\lambda_{\max} = \frac{hc}{5kT}$ .

$$\lambda_{\max} = \frac{6.626 \times 10^{-34} \text{ J s} \times 2.998 \times 10^8 \text{ m s}^{-1}}{5 \times 1.381 \times 10^{-23} \text{ J K}^{-1} \times 850. \text{ K}} = 3.38 \times 10^{-6} \text{ m at } 850. \text{ K. } \lambda_{\max} \text{ for } 1300. \text{ K and } 5500. \text{ K is}$$

$2.21 \times 10^{-6} \text{ m}$  and  $5.23 \times 10^{-7} \text{ m}$ , respectively.

**P12.15)** A beam of electrons with a speed of  $4.75 \times 10^4 \text{ m/s}$  is incident on a slit of width  $235 \text{ nm}$ . The distance to the detector plane is chosen such that the distance between the central maximum of the diffraction pattern and the first diffraction minimum is  $0.375 \text{ cm}$ . How far is the detector plane from the slit?

The diffraction minima satisfy the condition  $\sin \theta = \frac{n\lambda}{a}$ ,  $n = \pm 1, \pm 2, \dots$  and the first minimum is at

$\sin \theta = \pm \frac{\lambda}{a}$ . We choose the plus sign (the minus sign gives the distance from the slit in the opposite

direction) giving  $\sin \theta = \frac{\lambda}{a} = \frac{h}{mva} = \frac{6.626 \times 10^{-34} \text{ J s}}{9.109 \times 10^{-31} \text{ kg} \times 4.75 \times 10^4 \text{ m s}^{-1} \times 235 \times 10^{-9} \text{ m}} = 0.06517$   
 $\theta = 5.74 \text{ degrees}$

The distance  $d$  from the screen and the position of the first minimum  $s$  are related by

$$d = \frac{s}{\tan \theta} = \frac{0.375 \text{ cm}}{0.0652} = 5.74 \text{ cm}$$

**P12.18)** X-rays can be generated by accelerating electrons in a vacuum and letting them impact on atoms in a metal surface. If the 1525-eV kinetic energy of the electrons is completely converted to the photon energy, what is the wavelength of the X-rays produced? If the electron current is  $2.25 \times 10^{-5}$  A, how many photons are produced per second?

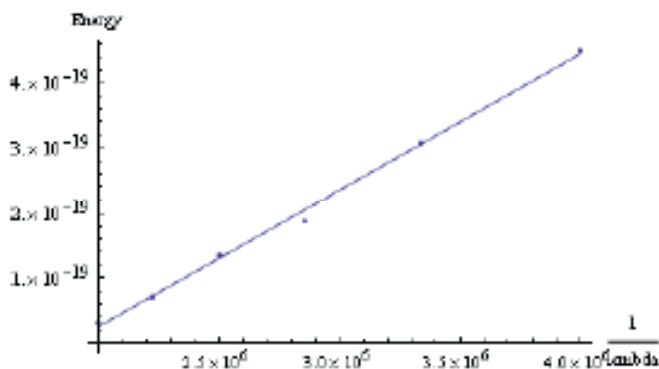
$$\lambda = \frac{hc}{E} = \frac{6.626 \times 10^{-34} \text{ J s} \times 2.998 \times 10^8 \text{ m s}^{-1}}{1525 \text{ eV} \times \frac{1.602 \times 10^{-19} \text{ J}}{\text{eV}}} = 0.813 \text{ nm}$$

$$n = \frac{\text{current}}{\text{charge per electron}} = \frac{2.25 \times 10^{-5} \text{ C s}^{-1}}{1.602 \times 10^{-19} \text{ C}} = 1.40 \times 10^{14} \text{ s}^{-1}$$

**P12.19)** The following data were observed in an experiment on the photoelectric effect from potassium:

<b><math>10^{19}</math> Kinetic Energy (J)</b>	4.49	3.09	1.89	1.34	0.700	0.311
<b>Wavelength (nm)</b>	250.	300.	350.	400.	450.	500.

Graphically evaluate these data to obtain values for the work function and Planck's constant.



The best-fit line is given by  $E(J) = -3.95831 \times 10^{-19} + 2.1032 \times 10^{-25} \frac{1}{\lambda}$ .

Because the slope is  $hc$ ,  $h = \frac{2.11171 \times 10^{-25} \text{ J m}}{2.998 \times 10^8 \text{ m s}^{-1}} = 7.02 \times 10^{-34} \text{ J s}$ . The work function is given by the

intercept of the line with the x axis at  $y = 0$ .  $\phi = \frac{hc}{\lambda_0} - E$  where  $\lambda_0 = \frac{2.1171 \times 10^{-25} \text{ J m}}{3.97362 \times 10^{-19} \text{ J}} = 5.31 \times 10^{-7} \text{ m}$

This gives  $\phi = 3.74 \times 10^{-19} \text{ J}$  or 2.33eV.

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**P12.22)** Assume that water absorbs light of wavelength  $3.50 \times 10^{-6}$  m with 100% efficiency. How many photons are required to heat 2.50 g of water by 1.00 K? The heat capacity of water is  $75.3 \text{ J mol}^{-1} \text{ K}^{-1}$ .

$$E = N h \nu = N \frac{h c}{\lambda} = n C_{p,m} \Delta T$$

$$N = \frac{m C_{p,m} \Delta T \lambda}{M h c} = \frac{2.50 \text{ g}}{18.02 \text{ g mol}^{-1}} \frac{75.3 \text{ J K}^{-1} \text{ mol}^{-1} \times 1.00 \text{ K} \times 3.50 \times 10^{-6} \text{ m}}{6.626 \times 10^{-34} \text{ J s} \times 2.998 \times 10^8 \text{ m s}^{-1}}$$
$$= 1.84 \times 10^{20}$$

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